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Late Cretaceous Gangdese intrusions of adakitic geochemical characteristics, SE Tibet: Petrogenesis and tectonic implications

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ABSTRACT

The Gangdese batholith emplaced from the Cretaceous to Eocene in southern Tibet has been widely regarded as the major constituent of an Andean-type convergent margin resulting from northward subduction of the Neo-Tethyan oceanic lithosphere under Asia. While the Gangdese batholith consists predominantly of calcalkaline rocks, we identify from the eastern part of the batholith a suite of epidote-bearing granodiorites that shows adakitic geochemical characteristics, marked with apparently higher La/Yb and Sr/Y, and lower Y and HREE, than other Gangdese rocks and common arc magmas. SHRIMP zircon U-Pb analyses of two of the samples yielded $^{206}\text{Pb}/^{238}\text{U}$ dates of 80.4 ± 1.1 and 82.7 ± 1.6 (2 σ) Ma, which constrain the emplacement ages of the adakitic rocks. Trace element modeling suggests that these rocks originated from partial melting of a garnet amphibolite source that, on the basis of the Sr and Nd isotope data [$I_{Sr}=0.7044-0.7048$; $\varepsilon_{Nd}(T)=+3.2$ to +0.9], we interpret to be a newly underplated, mafic lower crust, rather than the subducted Neo-Tethyan oceanic crust. This juvenile crust was produced by Cretaceous basaltic underplating above the mantle wedge and then thickened by the tectonic contraction owing to flattening of the Neo-Tethyan subduction, a process that also led to the adakitic magmatism. Our interpretation involving a Late Cretaceous flat-slab subduction and related orogenesis in southern Tibet is consistent with petrographic data, such as the occurrence of magmatic epidote and muscovite rimmed with resorption texture in the granodiorites, which indicate deepseated emplacement followed by rapid tectonic exhumation.

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1. Introduction

In southern Tibet, the elongated belt of granitoids that occurs in the Lhasa terrane north of the Yarlu–Zangpo suture zone (Fig. 1) has been termed the "Trans-Himalayan Batholith" (cf. Searle et al., 1987; and references therein) and regarded as a principal component of the Andean-type continental margin along South Asia resulting from northward subduction of the Neo-Tethyan oceanic lithosphere before the India–Asia collision (Allegrè et al., 1984; Yin and Harrison, 2000). Previous studies (e.g., Debon et al., 1986; Searle et al., 1987; Harris et al., 1988; TBGMR, 1993; Mo et al., 2005) indicated that the granitoids consist actually of a wide range of compositions from gabbro to granite, and the most abundant rock type is hornblende-biotite-bearing diorite and granodiorite that are geochemically comparable to the I-type calc-alkaline granitoids emplaced in the circum-Pacific regions (Chappell and White, 1974; Pitcher, 1982). As part of our study

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of the Trans-Himalayan magmatism, we describe in this paper a recently discovered suite of epidote-bearing granodiorites, characterized by adakitic geochemical affinities, from the eastern part of the Gangdese batholith, one of the largest intrusive bodies in southern Tibet. We report SHRIMP (sensitive high resolution ion microprobe) zircon U–Pb ages, mineralogical, major and trace element, and Sr–Nd isotope data for these rocks that suggest a magmatic origin by partial melting of a newly underplated and thickened mafic lower crust above the mantle wedge at ca. 80 Ma, when the Neo-Tethyan subduction was flattened. This study bears important implications for not only the petrogenesis of subduction-related adakitic magmatism in general, but also the precollisional tectonomagmatic evolution in southern Tibet where "collision-type" adakites were emplaced from the Late Oligocene to Middle Miocene after India started colliding with Asia (e.g., Chung et al., 2003; Hou et al., 2004; Guo et al., 2007).

2. Background

Tibet is essentially composed of four continental blocks or terranes that, from north to south, are the Songpan–Ganze, Qiangtang, Lhasa



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Fig. 1. Simplified geologic map showing principal igneous outcrops in the Lhasa terrane, southern Tibet (modified after Chung et al., 2003). Inset (a) denotes Cenozoic tectonic framework and distribution of Trans-Himalayan magmatic belt in SE Asia; (b) denotes sample localities, with stars indicating two samples dated by this study at 80 and 83 Ma. BNS – Bangong–Nujiang suture; YTS – Yarlu–Tsangpo suture; STDS – South Tibet detachment system.

and Tethyan Himalaya, separated by the Jinsha, Bangong–Nujiang and Yarlu–Tsangpo suture zones representing Paleo-, Meso- and Neo-Tethyan oceanic relicts, respectively (cf. Yin and Harrison, 2000). The Lhasa terrane, which dispersed from Gondwana and then drifted northward and finally amalgamated with the Qiangtang terrane in the Early Cretaceous (Kapp et al., 2005), consists primarily of Paleozoic to Paleogene sedimentary strata and associated igneous rocks (Yin and Harrison, 2000). The latter include the Linzizong volcanic successions and the widespread Trans-Himalayan intrusive bodies that have been divided into two suites (Fig. 1): a southern Gangdese belt dominated by dioritic and I-type compositions (Debon et al., 1986; Harris et al., 1988; TBGMR, 1993; Wen et al., 2003, 2004) and a northern plutonic belt comprised largely of peraluminous, or S-type, granitoids (Xu et al., 1985; Harris et al., 1988, 1990).

The duration of the Gangdese magmatism was previously constrained using Ar-Ar (e.g., Maluski et al., 1982; Copeland et al., 1995; Harrison et al., 2000), Rb-Sr (Debon et al., 1982; Jin and Xu, 1984; Xu et al., 1985) and zircon U-Pb age data (Schärer et al., 1984; Xu et al., 1985; Quidelleur et al., 1997; Mo et al., 2005). Most recently, based on a more detailed investigation of the Gangdese batholith, Wen et al. (in press) reported new SHRIMP zircon U-Pb ages of 25 samples that, together with the literature data, suggest two distinct stages of Gangdese magmatism in the Late Cretaceous (ca. 103-80 Ma) and early Paleogene (ca. 68-43 Ma), with a magmatic gap or quiescent period in between. Additionally, Chu et al. (2006) identified an early Jurassic, previously uncomprehended stage of Gangdese intrusion that implies that Neo-Tethyan subduction played a long-lasting role in the tectonic evolution of southern Tibet prior to the closure of Neo-Tethys Ocean and the start of the India-Asia collision. In this study, we document the emplacement age and geochemical characteristics of a suite of Late Cretaceous adakitic intrusion in the eastern part of the Gangdese batholith and explore its petrogenesis and significance to the regional tectonics.

3. Occurrence and samples

The studied area is located between Langxian and Lilong in the eastern Gangdese batholith (Fig. 1). Here intrusive rocks generally show median-grained and equigranular texture, with remarkably different mineral assemblages from other Gangdese granitoids. They are epidote–muscovite-bearing, consist mainly of plagioclase, K-feldspar, quartz and biotite, but lack hornblende and/or pyroxene that are common in the calc-alkaline Gangdese rocks. In some areas the intrusion is cut and deformed by the Renbu–Zedong Thrust, which Quidelleur et al. (1997) reported thermochronological data for suggesting Miocene thrusting activity. Quidelleur et al. (1997) also conducted ion microprobe analyses of zircon separates from the hanging wall granitoids that gave ²⁰⁶Pb/²³⁸U ages dominantly between ca. 80 and 100 Ma, indicating that the intrusion took place around the Late Cretaceous.

In this study, nine granodiorite samples were collected along the main road between Langxian and Lilong (Fig. 1). They are generally fresh, without visible alteration, and medium- to coarse-grained rocks, composed mainly of plagioclase, K-feldspar, quartz and biotite. Accessory phases include apatite, zircon, epidote, titanite, muscovite, Fe–Ti oxides, allanite and rutile. Two samples, T024 and T027, from the eastern and western parts of the sampling sites (Fig. 1), were subjected to zircon separation for SHRIMP U–Pb dating. We note that sample T027 was actually collected from a locality within the area studied by Quidelleur et al. (1997).

4. Analytical methods

Zircons were separated by heavy-liquid and magnetic methods. Insitu zircon U–Pb dating was carried out using the SHRIMP II at the Beijing SHRIMP Lab, the Chinese Academy of Geological Sciences, with analytical conditions following those reported in Liu et al. (2006) and Wen et al. (in press). U–Th–Pb isotope ratios were measured relative to the zircon standard TEMORA (Black et al., 2003).

Electron microprobe analyses on minerals were performed on polished thin sections at the Institute of Earth Sciences, Academia Sinica, Taipei, using a JXA®-8900R electron probe microanalyzer. Analyses were performed in wavelength dispersive mode with 15 keV accelerating voltage, 5 nA beam current and 1 µm spot diameter. More analytical conditions have been reported in Xu et al. (2004).

Powdered samples were analyzed for major elements by X-ray fluorescence (XRF) using a Rigaku® RIX 2000 spectrometer at the Department of Geosciences, National Taiwan University. The analytical

Table 1	
SHRIMP zircon U-Pb data of the granodiorites from SE Til	oet

Sample	U	Th	Th/U	f ₂₀₆	²⁰⁶ Pb*/ ²³⁸ U	²⁰⁷ Pb*/ ²³⁵ U	²⁰⁶ Pb*/ ²³⁸ U age
	(ppm)	(ppm)		(%)	(±1 σ)	(±1 o)	(Ma $\pm 1 \sigma$)
T024							
1.1	121	96	0.80	2.80	0.01221 ± 0.00023	0.0425 ± 0.0093	78.2±1.5
2.1	789	400	0.51	0.10	0.01247±0.00017	0.0808 ± 0.0026	79.9±1.1
3.1	359	282	0.78	0.23	0.01248 ± 0.00018	0.0814 ± 0.0040	79.9±1.2
4.1	393	180	0.46	0.13	0.01268 ± 0.00018	0.0829 ± 0.0049	81.2±1.2
5.1	934	515	0.55	0.43	0.01245±0.00017	0.0783 ± 0.0030	79.7±1.1
6.1	193	121	0.63	2.65	0.01282 ± 0.00035	0.0490 ± 0.0136	82.1±2.2
7.1	1288	471	0.37	-0.02	0.01313 ± 0.00017	0.0889 ± 0.0019	84.1±1.1
8.1	340	249	0.73	0.47	0.01241 ± 0.00018	0.0749 ± 0.0055	79.5±1.2
9.1	156	65	0.42	0.29	0.01220 ± 0.00023	0.0774±0.0125	78.2±1.5
10.1	452	351	0.78	-0.11	0.01282 ± 0.00018	0.0856 ± 0.0026	82.1±1.1
11.1	674	364	0.54	-0.02	0.01232 ± 0.00017	0.0811 ± 0.0024	78.9±1.1
12.1	443	366	0.83	0.51	0.01249 ± 0.00018	0.0795 ± 0.0044	80.0±1.1
Weighted mean (2 σ)							80.4±1.1
T027							
1.1	71	40	0.56	4.32	0.01272 ± 0.00043	0.0688 ± 0.0096	81.5±3.3
4.1	21	12	0.56	18.41	0.01314 ± 0.00060	0.0161 ± 0.0647	84.2±6.4
5.1	151	106	0.70	3.21	0.01280 ± 0.00038	0.0851 ± 0.0102	82.0±2.9
6.1	184	119	0.65	2.70	0.01275 ± 0.00038	0.0810 ± 0.0065	81.7±2.9
7.1	158	85	0.54	1.99	0.01342 ± 0.00039	0.0855 ± 0.0060	85.9±2.8
9.1	364	342	0.94	1.85	0.01266±0.00035	0.0717 ± 0.0050	81.1±2.7
11.1	63	39	0.62	7.72	0.01229 ± 0.00045	0.0627±0.0119	78.8±3.7
12.1	279	182	0.65	1.93	0.01315 ± 0.00037	0.0759 ± 0.0061	84.2±2.7
13.1	123	75	0.61	3.65	0.01345 ± 0.00042	0.0626±0.0088	86.1±3.2
15.1	29	19	0.64	10.62	0.01152 ± 0.00065	0.0487 ± 0.0209	73.8±5.9
17.1	64	35	0.55	6.71	0.01285 ± 0.00046	0.0610 ± 0.0177	82.3±3.8
18.1	39	23	0.58	14.13	0.01233±0.00047	0.0174±0.0294	79.0±4.4
20.1	110	70	0.63	4.85	0.01274±0.00039	0.0683 ± 0.0075	85.0±2.6
21.1	42	23	0.54	14.74	0.01188 ± 0.00045		76.9±6.9
22.1	272	214	0.79	1.45	0.01347 ± 0.00036	0.0804 ± 0.0040	86.2±2.8
23.1	130	161	1.23	5.06	0.01348 ± 0.00042	0.0635 ± 0.0159	86.3±3.7
Weighted mean (2 σ)							82.7±1.6
2.1	391	117	0.30	0.33	0.05385 ± 0.00140	0.3910±0.0117	338.1±9.0
3.1	469	352	0.75	-2.23	0.04157 ± 0.00108	0.4188 ± 0.0126	262.5±7.4
8.1	886	821.98	0.93	0.76	0.01654 ± 0.00043	0.0965 ± 0.0039	105.8±3.3
10.1	2005.29	136	0.07	0.26	0.01818 ± 0.00045	0.1207±0.0036	116.1±3.0
14.1	200	145	0.73	2.48	0.01555 ± 0.00044	0.0856 ± 0.0060	99.5±3.3
16.1	193	93	0.48	1.72	0.01933±0.00054	0.1212±0.0085	123.4±3.8
19.1	139	85	0.61	3.16	0.01534 ± 0.00044	0.0932 ± 0.0075	98.1±3.4

uncertainties are generally better than 5% for all elements. Trace elements were measured by inductively coupled plasma mass spectrometry (ICP-MS) using an Agilent[®] 7500 s machine in the same Department. The analytical accuracy and precision are generally better than 3%. Analytical details for the major and trace element measurements may be found in Chung et al. (2003) and Yang et al. (2005).

Whole-rock Sr and Nd isotope ratios were determined using a Finnigan MAT-262® mass spectrometer operated in static multicollector mode at the Department of Earth Sciences, National Cheng-Kung University. Measured ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd ratios were normalized to ⁸⁶Sr/⁸⁸Sr=0.1194 and ¹⁴⁶Nd/¹⁴⁴Nd=0.7219, respectively, for mass fractionation correction. During the period of data acquisition, the mean ⁸⁷Sr/⁸⁶Sr ratio of NBS987 standard was 0.710261±17 (*n*=8) and mean ¹⁴³Nd/¹⁴⁴Nd ratio of La Jolla standard was 0.511830±12 (*n*=12). Detailed chemical procedures and instrumentation specifications have been reported by Liu et al. (2007).

5. Results

5.1. Zircon U-Pb ages

Similar to zircon separates from other Gangdese rocks (Chu, 2006; Chu et al., 2006; Wen et al., in press), zircons of this study are mostly euhedral, with long to short prismatic shapes with average crystal lengths of ~100–200 µm and length-to-width ratios up to 4:1. Most zircons are transparent, colourless to slightly brown, and display oscillatory zoning typical of magmatic growth. Zircons with rounded or ovoid shape are, however, occasionally present in sample T027.

Zircon U–Pb isotope data are listed in Table 1 and plotted in the concordia diagrams (Fig. 2). Owing to the youth and relatively low U abundance of the zircons dated, large uncertainties on ²⁰⁷Pb/²³⁵U ratios and calculated ages are observed. Therefore, weighted means of pooled ²⁰⁶Pb/²³⁸U ages are taken to represent the crystallization ages that are reported with uncertainties at two standard deviations (2σ) or 95% confidence level. A total of twelve analyses of zircons from sample T024 yielded rather uniform ²⁰⁶Pb/²³⁸U ratios and a weighted mean of ${}^{206}\text{Pb}/{}^{238}\text{U}$ ages at 80.4±1.1 Ma (Fig. 2a). The Th and U concentrations range from 515 to 65 and 1288 to 121 ppm, respectively, yielding Th/U ratios between 0.83 and 0.42 that are in good accordance with those of igneous zircons (Belousova et al., 2002). Although zircons from sample T027 are slightly more complex, 16 out of 23 analyses yielded a mean 206 Pb/ 238 U age at 82.7±1.6 Ma (Fig. 2b), with Th and U contents varying from 342 to 12 and 364 to 21 ppm, and Th/U ratios from 1.23 to 0.54. The above two U-Pb ages, statistically overlapped if considering the 2σ analytical uncertainties, delineate an important intrusive event at ca. 83-80 Ma in this part of the Gangdese batholith. Our results, moreover, support and refine the previously published zircon U-Pb age constraints that span between ca. 80 and 100 Ma with an age cluster at ca. 82-86 Ma (Quidelleur et al., 1997).



Fig. 2. Zircon U–Pb concordia diagrams of samples (a) T024 and (b) T027. Note that the mean ages shown are $^{206}\rm{Pb}/^{238}\rm{U}$ ages.

In sample T027, analyses of the remaining seven zircon grains yielded older 206 Pb/ 238 U ages between ca. 98 and 338 Ma (Table 1). With the exception of one grain that has apparently higher U content (2005 ppm) and thus low Th/U ratio (0.07), the other six grains exhibit Th/U ratios (0.93–0.30) in the range typical of igneous zircons, and are interpreted to be inherited zircons produced by earlier magmatism in the region.

5.2. Mineral chemistry

As mentioned above, the rocks are characterized by the presence of magmatic epidote and muscovite as accessory phases that, along with the coexisting feldspars, bear unique petrologic significance. The compositions of these phases are sensitive to the P–T conditions of the host magma during emplacement, and therefore were carefully analyzed with the electron microprobe. Representative results on sample T024 are given in Tables 2–4.

5.2.1. Magmatic epidote

Epidote, which is observed in all the samples, has an average grain size of ~1 mm, similar to that of the major mineral constituents (Fig. 3, a–b). The epidote, occurring mostly as euhedral grains adjacent to biotite, occasionally shows embayed, wormy contacts with plagioclase and quartz (Fig. 3c), and often contains inclusions of phases such as allanite and/or rutile and biotite (Fig. 3d). The crystal form, resorption texture and fresh appearance of the epidote are consistent with those

Table	2
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Representative electron microprobe data of epidote

Minerarl	Epidote	e (n=40)						
Analysis	T-3	T-4	T-7	T-8	T-13	T-14	Mean	σ
SiO ₂	38.61	37.39	38.73	37.70	38.28	38.02	38.06	0.44
TiO ₂		0.08	0.23	0.20	0.10	0.16	0.13	0.06
Al_2O_3	22.12	22.13	22.72	22.29	22.78	22.78	22.65	0.54
Cr_2O_3	0.03		0.05				0.01	0.02
Fe ₂ O ₃	12.08	12.97	12.64	12.07	11.89	12.78	12.27	0.62
MnO	0.46	0.51	0.59	0.64	0.58	0.56	0.55	0.08
MgO	0.04	0.03	0.05	0.02		0.05	0.02	0.02
NiO		0.01	0.22	0.11		0.01	0.03	0.05
CaO	22.34	23.14	22.80	23.21	22.72	22.86	23.06	0.48
Na ₂ O								0.01
K ₂ O		0.03	0.01		0.04		0.01	0.01
P_2O_5	0.02	0.03			0.02	0.03	0.02	0.03
F	0.57	0.01	0.36			0.12	0.06	0.13
Cl				0.05	0.01		0.01	0.01
Total	96.27	96.33	98.40	96.29	96.42	97.37	96.89	0.96
Number of	ions on th	ne basis of	25 oxygen	s				
Si	6.40	6.22	6.30	6.25	6.30	6.24	6.26	0.07
Ti		0.01	0.03	0.03	0.01	0.02	0.02	0.01
Al	4.32	4.34	4.35	4.36	4.42	4.41	4.39	0.08
Cr			0.01					
Fe	1.47	1.58	1.51	1.47	1.44	1.54	1.49	0.08
Mn	0.07	0.07	0.08	0.09	0.08	0.08	0.08	0.01
Mg	0.01	0.01	0.01			0.01		
Ni			0.03	0.01				
Ca	3.96	4.12	3.97	4.12	4.01	4.02	4.06	0.08
Na								
K P		0.01			0.01			
Ps	26	27	26	26	25	26	25.7	1.3

Analysis in wt.%; total Fe measured as Fe₂O₃.

Ps: Molar Pistacite $[Fe^{3+}/(Fe^{3+}+Al^{3+})\times 100]$.

ble	3			

Та

	Representative e	lectron	microprob	be data	of	muscov	ite
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Minerarl	Muscov	vite (n=5	1)					
Analysis	T1-23	T1-24	T1-41	T1-42	T2-01	T2-05	Mean	σ
SiO ₂	46.50	47.03	46.84	47.05	46.40	46.51	46.47	0.66
TiO ₂	1.06	1.17	1.22	0.56	0.92	0.80	0.71	0.24
Al ₂ O ₃	27.39	28.20	26.85	27.30	27.32	28.55	27.91	0.59
Cr ₂ O ₃			0.02					
Fe ₂ O ₃	6.21	5.84	5.84	4.65	6.93	5.85	5.87	0.55
MnO	0.03	0.10	0.15	0.05	0.09	0.04	0.07	0.05
MgO	2.09	2.01	2.25	2.11	2.26	1.95	2.15	0.20
NiO	0.20		0.09	0.17			0.01	0.04
CaO				0.03	0.03		0.02	0.02
Na ₂ O	0.25	0.24	0.20	0.21	0.25	0.26	0.24	0.04
K ₂ O	10.60	10.92	11.25	11.06	10.28	11.17	11.00	0.28
P ₂ O ₅			0.10	0.01				
F	0.26			0.04			0.11	0.15
Cl		0.01				0.03	0.01	0.01
Total	94.59	95.52	94.80	93.25	94.48	95.16	94.58	0.84
Number of ions on	the basis	of 11 oxy	gens					
Si	3.23	3.22	3.24	3.28	3.18	3.17	3.19	0.03
Ti	0.06	0.06	0.06	0.03	0.05	0.04	0.04	0.01
Al	2.24	2.27	2.19	2.24	2.21	2.29	2.26	0.04
Cr								
Fe	0.36	0.33	0.34	0.27	0.36	0.30	0.31	0.03
Mn		0.01			0.01			
Mg	0.22	0.20	0.23	0.22	0.23	0.20	0.22	0.02
Ni	0.01		0.01	0.01				
Ca								
Na	0.03	0.03	0.03	0.03	0.03	0.03	0.03	0.01
К	0.94	0.95	0.99	0.98	0.90	0.97	0.96	0.03
Р			0.01					
P(kbar) at 765 °C	9.6	9.3	10.0	11.0	8.3	8.0	8.5	0.9

Analysis in wt.%; total Fe measured as Fe₂O₃.

Crystallization pressure (P) calculated using the equation of Anderson (1996).

Table 4		
Representative electron	microprobe data	of feldspar

Minerarl	Plagiclase	(oligoclase)					Plagiclase	(albite)			Orthocla	ise	
Analysis	T2-12	T2-45	T2-46	T1b-76	T2b-41	T2b-42	T2-43	T2-44	T1b-74	T2b-47	T2-1	T2-2	T2-5
SiO ₂	61.54	63.10	62.32	61.52	61.59	61.72	68.40	68.76	68.57	67.62	64.83	63.39	66.41
TiO ₂	0.04	0.06	0.02	0.01	0.08		0.03			0.09	0.08		0.02
Al_2O_3	23.79	22.66	23.07	23.59	23.68	23.46	19.86	19.66	19.47	20.14	18.19	18.28	18.15
Cr_2O_3								0.02			0.01		
Fe ₂ O ₃	0.09			0.02	0.05	0.15			0.19	0.13		0.02	
MnO						0.02	0.01			0.02	0.02		
MgO		0.02	0.02										
NiO	0.06	0.10						0.11	0.05	0.05	0.03	0.01	
CaO	5.72	5.11	5.00	5.83	5.75	5.87	0.96	0.32	0.55	0.85		0.01	
Na ₂ O	8.41	8.99	8.41	8.27	8.78	8.84	11.03	11.35	11.11	11.69	0.54	0.44	0.60
K ₂ O	0.28	0.07	0.09	0.26	0.17	0.25	0.07	0.06	0.09	0.05	16.13	15.90	16.18
P_2O_5	0.07	0.01	0.04					0.01			0.05		0.08
F			0.12	0.05		0.15	0.15	0.07	0.11	0.27			
Cl	0.01		0.03							0.01	0.02		
Total	100.00	100.13	99.11	99.56	100.10	100.45	100.51	100.36	100.14	100.87	99.90	98.06	101.44
Number of io	ns on the basi	s of 32 oxyger	15										
Si	10.95	11.17	11.13	8.58	10.95	10.97	11.91	11.97	11.98	11.80	12.00	11.95	12.08
Ti	0.01	0.01			0.01					0.01	0.01		
Al	4.99	4.73	4.86	3.88	4.96	4.91	4.08	4.03	4.01	4.14	3.97	4.06	3.89
Cr													
Fe	0.01				0.01	0.02			0.03	0.02			
Mn													
Mg		0.01											
Ni	0.01	0.01						0.02	0.01		0.01		
Ca	1.09	0.97	0.96	0.87	1.09	1.12	0.18	0.06	0.10	0.16			
Na	2.90	3.08	2.91	2.24	3.03	3.04	3.72	3.83	3.76	3.95	0.19	0.16	0.21
К	0.06	0.02	0.02	0.05	0.04	0.06	0.02	0.01	0.02	0.01	3.81	3.82	3.75
Р	0.01		0.01								0.01		0.01
Ab (mol%)	71.6	75.8	74.9	70.9	72.8	72.2	95.0	98.1	96.8	95.9	4.8	4.1	5.3
An (mol%)	26.9	23.8	24.6	27.6	26.3	26.5	4.6	1.5	2.7	3.8	0.0	0.1	0.0
Or (mol%)	1.6	0.4	0.5	1.4	0.9	1.3	0.4	0.3	0.5	0.3	95.2	95.9	94.7

Analysis in wt.%; total Fe measured as Fe₂O₃.

of primary epidote of magmatic origin (cf. Zen and Hammastrom, 1984; Schmidt and Poli, 2004). In addition, the edipotes we analyzed have low TiO₂ (<0.2 wt.%) and restricted pistacite (Ps) contents, i.e., atomic Fe³⁺/(Fe³⁺+Al³⁺), between 23% and 28% (Fig. 4), that are compatible with primary epidotes reported in the South American (Sial et al., 1999) and North American Cordilleras (Zen and Hammastrom, 1984; Barth, 1990; Brandon et al., 1996), and match the chemical criteria of magmatic epidotes (Zen and Hammastrom, 1984; Schmidt and Poli, 2004).

5.3. Celadonitic muscovite

Muscovite, with a grain size from ~0.3 to >1 mm in length (Fig. 3a), often occurs in association with epidote as an euhedral phase that occasionally is resorbed by quartzo-feldspathic matrix (Fig. 3c). Microprobe analyses (Table 3) indicate that this phase has high TiO₂ content (from ~0.3 to 1.2 wt.%), typical of magmatic muscovite (>0.4 wt.%; Anderson, 1996). The high celadonitic components (i.e., the amount of Mg and Fe that substitutes for Al in the octahedral site, and Si that substitutes for Al in the tetrahedral site) observed here are also typical of primary muscovite in peraluminous granitoids reported elsewhere (Barth, 1990; Anderson, 1996).

5.4. Feldspar

Plagioclase and K-feldspar are common, with the former being the most abundant, in the rocks (Fig. 3, a–b). While K-feldspar of orthoclase composition ($Or \approx 95$) is typical, there are two generations of plagioclase composed of oligoclase (An_{24} – An_{28}) and albite (An_{2-} – An_{5})(Table 4). Petrographic evidence shows that oligoclase is the main crystalline phase, while albite crystallized along with quartz from

incongruent melts of the oligoclase's margins at a later stage (Fig. 3e). In addition, the late-stage phase of quartz occasionally occurs as droplets along the boundaries of orthoclase and muscovite (Fig. 3f). These disequilibrium textures, together with the above-described resorption features rimming epidote and muscovite, imply that the minerals crystallized under unstable conditions at depth, and then underwent decompressional melting along crystal margins owing to rapid upward transport (Nelson and Montana, 1992).

5.5. Major and trace elements

Whole-rock major and trace element data are presented in Table 5. The rocks show uniform major element compositions, with SiO₂ contents of ~65–70 wt.% and Al₂O₃ of ~16–18 wt.%. With the exception of sample T212 (which has higher K₂O of ~3 wt.%), these granodiorites are medium-K in nature, with Na₂O>4.1 wt.% and Na₂O/K₂O>2. In comparison with other Gangdese rocks that range from gabbro to granite, they all are peraluminous, with A/CNK values, i.e., molar Al₂O₃/(CaO+Na₂O+K₂O), >1.0 (Fig. 5) and have low MgO (<1.0 wt.%) or Mg# (~37–31).

More interestingly, the rocks exhibit strongly fractionated rare earth element (REE) patterns (Fig. 6) characterized by high $(La/Yb)_N$ ratios (~42–9; Table 5) and low heavy REE (HREE; e.g., Yb=0.4–1.1 ppm) and Y contents (~4–12 ppm) that are distinctly different from other Gangdese rocks or the "normal" calc-alkaline arc lavas (Figs. 6 and 7). In addition, these rocks contain high Sr contents (~780–620 ppm), leading to high Sr/Y ratios (~200–60; Table 5). All such elemental characteristics, in particular elevated Sr/Y and La/Yb ratios (Fig. 8), are discriminating features observed in adakites from the circum-Pacific subduction zones (Defant and Drummond, 1990; Martin, 1999). Hence, we conclude that the peraluminous



Fig. 3. Photos (a, b, crossed polarized light) and backscattered electron images (c–f) showing textures and mineral assemblages of the granodiorites. (a) Euhedral epidote (Ep) and muscovite (Mus) showing resorption texture surrounded by plagioclase (Pl), orthoclase (Or) and quartz (Qtz), (b) corroded epidote with allanite (Al) core rimmed by biotite (Bt), (c) euhedral muscovite and epidote, with wormy quartz, (d) embayed, vermicular texture of epidote in contact with quartz, plagioclase, and rutile (Ru)/biotite (Bt) inclusions, (e) resorbed epidote showing two stages of dissolution, in which the early absorbing orthoclase is resorbed by oligoclase with an albite rim, (f) corroded muscovite in irregular contact with orthoclase and plagioclase rimmed by albite, arrows denote droplets of quartz that form along boundaries of muscovite and orthoclase.

granodiorites of this study are marked by adakitic geochemical compositions, and refer to them hereafter as "the Gangdese adakites".

5.6. Sr-Nd isotopes

Sr and Nd isotope results of the Gangdese adakites are also listed in Table 5. As a whole, they show slightly lower $\varepsilon_{Nd}(T)$ values (+3.0 to

+0.9) and higher initial 87 Sr/ 86 Sr ratios (~0.7044 to 0.7048) than the Cretaceous Gangdese calc-alkaline granitoids that are characterized by juvenile mantle isotopic signatures (Fig. 9). This implies that the adakitic magma generation involved a slightly higher extent of assimilation of material of the continental crust, relative to the associated calc-alkaline magmas. However, in light of their positive $\varepsilon_{Nd}(T)$ values, and thus younger Nd model ages or average crustal



Fig. 4. Histogram of pistacite (Ps) contents of epidote. Comparison data of magmatic and metamorphic epidotes are from Zen and Hammastrom (1984), Barth (1990), Brandon et al. (1996) and Sial et al. (1999).

residence time (i.e., T_{DM} =0.51–0.80 Ga; Table 5), the juvenile mantle component must have also played an important role in the petrogenesis of the Gangdese adakites.

6. Discussion

6.1. Magma source – Sr and Nd isotopic constraints

In subduction zones, as exemplifed by the Andes (Kay and Kay, 2002), there are three ways to form adakitic magmas given the consensus that a mafic source containing garnet and/or amphibole as the residual phases is required (cf. Rapp et al., 1991). The magma generation involves three candidate sources that are subducting oceanic crust (Defant and Drummond, 1990; Martin, 1999), underplated and thickened basaltic lower crust above the mantle wedge (Atherton and Petford, 1993; Kay and Kay, 1993), and subductioneroded forearc continental crust (Kay and Mpodozis, 2001). These distinct sources may be delineated using isotope compositions of the adakitic magmas.

As illustrated in Fig. 9, Sr and Nd isotope ratios of the Gangdese adakites preclude them from any direct genetic links with the Neo-Tethyan oceanic crust and subduction erosion related processes. Moreover, the low Mg# values (~31-37) and low Ni (0.4-5.8 ppm) and Cr (<11.4 ppm) contents of the Gangdese adakites differ markedly from those of slab-derived adakitic melts (Mg#>~47; Martin, 1999; Smithies, 2000) that would normally interact with the mantle wedge peridotite during magma ascent. Note that in Fig. 9 we assume the isotopic composition of the subducted oceanic slab to be similar to that of the Neo-Tethyan ophiolites from the Yarlu Tsangpo suture zone (Xu and Castillo, 2004; Zhang et al., 2005). These ophiolitic rocks have high positive $\varepsilon_{Nd}(T)$ values (\geq +8, except one analysis around +6), indicating the depleted mantle nature within the Neo-Tethyan asthenosphere (Xu and Castillo, 2004; Zhang et al., 2005). If melting occurred during subduction of this kind of oceanic crust, the melts generated would require a large amount of mixing with the continental crust (>15-20%) to form the Sr-Nd isotope ratios observed in the Gangdese adakites. This is considered very unlikely because such high degrees of crustal contamination should significantly change the compositions of the slab melts and thus their trace element characteristics (e.g., the high La/Yb and Sr/Y ratios) would become indistinct or even disappear.

The Gangdese adakites, in fact, show a resemblance to the associated calc-alkaline rocks in terms of Sr-Nd isotope compositions. This isotopic similarity strongly suggests a common magma source between the Cretaceous Gangdese arc rocks and later adakites. Formation of large batholiths like the ones in the Andes (e.g., Petford and Atherton, 1996) or California (e.g., Ducea, 2001), and the Gangdese (Chu, 2006; Wen, 2007), has generally been attributed to melting of hydrous, underplated basaltic rocks at lower crust levels that may best account for the voluminous silicic rocks formed. This scenario is actually in good consistency with the isotope data (Fig. 9). As shown by the mixing calculation, the Gangdese adakites can be generated by melting of a Gangdese gabbro representing the mafic lower crust, with $\varepsilon_{\rm Nd}(T)$ value of +3.7 and ${}^{87}{\rm Sr}/{}^{86}{\rm Sr}_i$ of 0.70435, accompanied by small degrees (<2%, or ~5% at most) of contamination by the upper continental crust during magma ascent. We therefore propose that the Gangdese adakites were derived from melting of a newly underplated, and then thickened, basaltic lower crust above the mantle wedge in the Neo-Tethyan subduction zone. Considering that there are three grains of detrital zircons dated at ca. 100 Ma in sample T027 (Table 1), broadly coeval to the emplacement age of a granite body of ca. 103 Ma near Langxian (Fig. 1; data from Wen et al., in press), we further argue the basaltic magma underplating to have taken place since ca. 100 Ma or the early Late Cretaceous associated with the Gangdese calc-alkaline magmatism. Subsequent thickening and partial melting of the crust to form the adakitic intrusion are ascribed to a flat-slab stage of the Neo-Tethyan subduction (see below).

6.2. Trace element modeling of magma generation

Generation of the Gangdese adakites by melting the newly underplated, thickened mafic lower crust, which could be of eclogite or garnet amphibolite composition (Rapp et al., 1991; Chung et al., 2003), may be modeled using specific element pairs such as La/Yb and Sr/Y (Fig. 8). With garnet being a residual phase in the melting regime, the magma generation must have occurred at high-pressure or deep crustal levels (\geq 1.2 GPa or 40 km; Rapp et al., 1991). Considering the mildly flattened HREE patterns observed in most samples (Fig. 6), which implies amphibole to have played a more important role than garnet during partial melting, garnet amphibolite rather than eclogite is more likely to be the source lithology.

In the La/Yb vs. Yb modeling calculation (Fig. 8b), the composition of a Gangdese gabbro (i.e., sample T036B, used also for the Sr–Nd mixing calculation) was utilized as the "starting material" representative of basaltic underplate that transformed to a garnet-bearing amphibolite in the thickened lower crust before partial melting. The modeling results show that the Gangdese adakites can be generated by ~10–20% melting of the presumed garnet amphibolite with its garnet constituents varying from ~3 to 15% (Fig. 8b). In this modeling the (La/Yb)_N and Yb_N correlation seems better interpreted in terms of combined variations in both the garnet constituents and the melting degrees. Note that eclogite melting was also modeled and the result does not match with the data (Fig. 8b).

6.3. Deep-seated emplacement followed by rapid exhumation

It has been well documented that magmatic epidote crystallizes in granitic magmas only at pressure $\geq 0.6-0.8$ GPa (Zen and Hammastrom, 1984; Schmidt and Poli, 2004). Plutonic rocks bearing magmatic epidote, hence, must have formed under even higher pressure that corresponds to lower crustal levels. Celadonitic muscovite is also sensitive to the pressure of crystallization, and thus can be used to estimate the depth of magma emplacement using phengite geobarometry (Massonne and Schreyer, 1987; Anderson, 1996) in combination with zircon saturation thermometry (Watson and Harrison, 1983). Accordingly, we calculated the emplacement P–T conditions for the

Table 5	;
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Geochemical data of the ca. 80 Ma Gangdese granitoids from Southern Tibet

Locality Impsian	Sample	T212	T027	T213	T215	T026	T216A	T217	T024	T218B
Imagemende Latinular 93.37 E 23.00° N 93.47 E 24.00° N 93.47 E 24.00° N 93.47 E 24.00° N 93.47 E 24.00	Locality	Langxian	Langxian	Langxian	Langxian	Langxian	Lilong	Lilong	Lilong	Lilong
Jatimade 2800° N 2800° N 2800° N 2800° N 2812° N <	Longtitude	93.31° E	93.32° E	93.34° E	93.41° E	93.44° E	93.61° E	93.64° E	93.75° E	93.75° E
Attribute (m) 3059 3074 3039 3014 3005 3023 3022 3065 Aper (Ma) R27 ± 15 R27 ± 15 R04 ± 1.1	Latitude	29.00° N	29.00° N	29.04° N	29.10° N	29.12° N	29.17° N	29.14° N	29.14° N	29.14° N
Age (Ma) B2.7±1.6 B0.4±11 (wt.3) 55.0 65.07 65.28 70.11 67.28 68.33 67.79 (wt.3) 0.66 0.77 0.06 0.70 0.72 0.22 (wt.3) 0.67 0.68 0.70 0.73 0.22 0.25 (wt.3) 0.68 0.70 0.70 0.22 0.61 0.06 0.07 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.67 0.66 0.63 0.63 0.63 0.63 0.63 0.63 0.63 0.63 0.61 0.20 0.61 0.20 0.61 0.20 0.61 0.20 0.61 0.20 0.61 0.20 0.61 0.20 <th>Altitude (m)</th> <th>3059</th> <th>3092</th> <th>3074</th> <th>3039</th> <th>3014</th> <th>3006</th> <th>3023</th> <th>3022</th> <th>3065</th>	Altitude (m)	3059	3092	3074	3039	3014	3006	3023	3022	3065
pmx3 solution solution <th< th=""><th>Age (Ma)</th><th></th><th>82.7±1.6</th><th></th><th></th><th></th><th></th><th></th><th>80.4±1.1</th><th></th></th<>	Age (Ma)		82.7±1.6						80.4±1.1	
SiD2 65.33 66.45 65.07 68.24 70.11 67.29 68.78 69.73 72.20 Ab,O, 16.56 16.57 18.04 17.20 15.55 17.04 17.79 17.65 16.56 KCO* 31.6 2.87 18.34 2.14 1.14 2.22 18.3 2.22 2.50 MO 0.07 0.09 0.11 0.05 0.04 0.35 0.34 0.34 0.34 0.34 0.35 0.35 0.34 0.34 0.34 0.34 0.34 0.35 0.34 0.34 0.34 0.34 0.34 0.35 0.35 0.35 0.35 0.35 0.35 0.35 0.35 0.35 0.35 0.36 0.06 0.10 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00 0.06 0.00	(wt.%)									
TiO, 0.36 0.36 0.24 0.16 0.30 0.23 0.26 0.23 ReD 1165 1157 11814 1220 1555 17.04 17.35 17.85 1629 ReD 316 2.27 2.30 0.35 0.31 0.22 2.31 0.31 0.22 0.31 0.31 0.32 0.31 0.31 0.32 0.31 0.31 0.32 0.31 0.31 0.32 0.31 0.31 0.32 0.31 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.45 0.41 0.4 1.11 1.11 1.12 1.12 1.13 0.4 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 0.44 </td <td>SiO₂</td> <td>65.93</td> <td>66.45</td> <td>65.07</td> <td>68.28</td> <td>70.11</td> <td>67.89</td> <td>68.78</td> <td>69.53</td> <td>67.79</td>	SiO ₂	65.93	66.45	65.07	68.28	70.11	67.89	68.78	69.53	67.79
Abb. Bbb Bbb<	TiO ₂	0.36	0.36	0.36	0.24	0.16	0.30	0.23	0.26	0.29
rep 3.0 2.4 2.4 1.4 2.32 1.03 2.42 2.30 Ma00 955 8.3 1.00 0.49 9.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.44 0.32 0.01 0.33 0.54 0.34 0.32 0.01 0.33 0.35 0.35 0.35 0.01 0.00 0.00 0.01 0.00 0.00 0.01 0.00 0.01 0.00 0.01 0.01 0.00 0.01	Al ₂ O ₃	16.56	16.57	18.04	17.20	15.95	17.04	17.59	17.65	16.96
map 0.05 0.03 0.04 0.03 0.04 0.04 0.05 0.03 CaD 335 335 440 321 2.61 3444 323 3.54 3.43 NayO 368 413 418 4.48 4.38 4.46 4.46 4.41 4.26 KO 2.97 1.84 0.12 2.09 2.20 1.82 2.10 1.73 2.15 Pop . <	MpO	3.10	2.87	3.43 0.11	2.14	1.41	2.52	1.83	2.22	2.50
Ca0 3.35 3.83 4.40 2.21 2.61 3.44 3.22 3.54 3.45 Na_0 3.68 4.13 4.18 4.44 4.38 4.46 1.0 0.0	ΜσΩ	0.07	0.09	1.00	0.00	0.04	0.61	0.00	0.58	0.00
Na-0 36.8 4.13 4.18 4.48 4.38 4.46 4.46 4.41 4.26 P-O ₅ 0.13 0.14 0.14 0.06 0.04 0.08 0.06 0.00 0.09 Me* 37 36 37 31 31 32 32 34 35 Me* 37 36 37 31 31 32 32 34 35 Me* 37 36 10.1 1.09 1.11 1.11 1.19 31.7 34 7.2 34 7.2 34 7.2 34 7.2 7.4 7.4 7.4 7.4 7.4 7.4 7.3 7.3 7.7 7.4 7.4 7.4 7.3 7.7 7.4 7.4 7.3 7.7 7.4 7.4 7.3 7.7 7.4 7.4 7.5 7.4 7.5 7.4 7.7 7.4 7.5 7.4 7.7 7.5 7.5 7.6	CaO	3.35	3.83	4.40	3.21	2.61	3.44	3.32	3.54	3.45
K_0 2.97 1.84 2.12 2.09 2.20 1.82 2.10 1.73 2.15 Pop. 0.13 0.14 0.14 0.06 0.04 0.08 0.06 0.00 0.09 A(CNK 1.08 1.05 1.11 1.11 1.02 1.2 1.13 1.09 (pm) -	Na ₂ O	3.68	4.13	4.18	4.48	4.38	4.46	4.46	4.41	4.26
P ₀ 0 0.13 0.14 0.14 0.06 0.04 0.08 0.06 0.10 0.09 Mg# 37 36 37 31 31 32 32 34 33 (pm) Sc 10.2 7.3 9.4 10.1 3.0 6.9 9.9 6.1 7.2 V 4.4 3.24 n.d. n.d. 8.77 n.d. n.d. 11.40 12.5 S.86 2.81 12.5 Ca 37.0 58.5 15.1 8.4 15.5 5.5 8.6 2.81 12.5 Ca 37.0 13.2 13.7 13.3 35.9 49.9 40.9 12.2 <td>K₂O</td> <td>2.97</td> <td>1.84</td> <td>2.12</td> <td>2.09</td> <td>2.20</td> <td>1.82</td> <td>2.10</td> <td>1.73</td> <td>2.15</td>	K ₂ O	2.97	1.84	2.12	2.09	2.20	1.82	2.10	1.73	2.15
A(CNK 1.08 1.05 1.11 1.11 1.19 1.12 1.13 1.09 Mg# 37 36 37 31 31 32 32 34 33 GpmJ Sc 1.02 7.3 9.4 1.01 3.0 6.9 9.9 6.1 7.2 C 0.19 9.07 n.d. n.d. 4.45 n.d. n.d. n.d. 1.10 1.11 1.11 1.10 1.11 1.11 1.10 1.11<	P ₂ O ₅	0.13	0.14	0.14	0.06	0.04	0.08	0.06	0.10	0.09
Mg# 37 36 37 31 31 32 32 34 33 (ppm)	A/CNK	1.08	1.05	1.05	1.11	1.11	1.09	1.12	1.13	1.09
(ppm) Sc 10.2 7.3 9.4 10.1 3.0 6.9 9.9 6.1 7.2 V 4.4 32.4 n.d. n.d. 87.7 n.d. n.d. 11.40 n.d. Mn 506 62.5 72.8 462 360 400 410 52.0 457.7 Ni 1.6 5.2 1.1 0.4 1.2 0.4 0.7 5.8 0.5 Cu 37.0 558.5 15.1 8.4 1.15 5.5 8.6 28.1 2.4 Ca 18.7 7.3 18.5 17.9 16.0 17.8 18.2 10.6 7.7 5.3 3.5 9.9 40.9 40.2 46.7 4.4 3.0 5.0	Mg#	37	36	37	31	31	32	32	34	33
Sc 10.2 7.3 9.4 10.1 3.0 6.9 9.9 6.1 7.2 V 4.4 32.4 n.d. n.d. 8.4 n.d. n.d. 1.8 n.d. n.d. 1.8 n.d. 1.8 n.d. 1.6 2.5 1.9 3.1 2.7 Ni 1.6 5.2 1.1 0.4 1.2 0.4 0.7 5.8 0.5 Cu 37.0 55.8.5 15.1 8.4 1.15 5.5 8.6 2.8.1 1.2.5 Ca 1.8.7 1.73 18.5 17.9 16.0 1.7.8 1.8.2 19.6 1.7.7 Rb 90.3 1.16 6.8 9.1 4.1 7.1 5.0 5.0 Y 9.9 9.3 1.16 6.8 9.1 4.1 7.1 5.0 5.0 Y 9.9 9.3 1.6 8 9.1 4.1 0.77 2.9.0 2.17	(ppm)									
V 4.4 3.2.4 n.d. n.	Sc	10.2	7.3	9.4	10.1	3.0	6.9	9.9	6.1	7.2
Li u.u. y.u.y y.u.y u.u. u.u. <thu.u.< th=""> u.u. u.u. <thu< td=""><td>V</td><td>4.4</td><td>32.4</td><td>n.d.</td><td>n.d.</td><td>87.7</td><td>n.d.</td><td>n.d.</td><td>11.8</td><td>n.d.</td></thu<></thu.u.<>	V	4.4	32.4	n.d.	n.d.	87.7	n.d.	n.d.	11.8	n.d.
mm job d.2.1 j.2.6 i.0.2 i.0.3 i.0.3 <thi.0.3< th=""> i.0.3 i.0.3</thi.0.3<>	Cl' Mp	0.19	9.07	11.0. 729	11.0.	4.45	11.0.	11.0. /10	520	11.0. 457
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		200 41	025 4.4	720 4.4	20	16	400	410	31	437
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Ni	16	5.2	11	0.4	1.0	0.4	0.7	5.8	0.5
Zn 53.4 60.9 62.4 49.0 42.2 46.7 44.3 51.2 49.3 Ga 18.7 17.3 18.5 17.9 16.0 17.8 18.2 19.6 17.7 Rb 90.3 41.3 45.1 54.5 53.3 35.9 49.9 40.9 42.6 Sr 622 622 688 757 626 768 780 738 719 Y 9.9 9.3 11.6 6.8 9.1 4.1 7.1 5.0 5.0 Zr 120 106 102 137 13.07 2.97 3.46 40.5 2.79 Cs 3.87 1.32 1.73 2.86 1.52 0.82 1.41 0.77 0.82 Ba 492 407 479 5.4 49.2 3.7 3.1.3 2.73 2.1.7 Ce 36.3 3.5.4 2.75 3.5.4 49.2 3.7 3.1.3 2.73 2.1.7 Ce 36.3 3.5.4 2.79 2.61	Cu	37.0	558.5	15.1	8.4	11.5	5.5	8.6	28.1	12.5
Ga 18.7 17.3 18.5 17.9 16.0 17.8 18.2 19.6 17.7 Rb 90.3 41.3 45.1 54.5 53.3 55.9 49.9 40.9 42.6 Sr 622 622 622 688 757 626 768 780 738 719 Y 9.9 9.3 116 6.8 9.1 4.1 71 5.0 5.0 Zr 120 106 102 137 135 139 122 139 117 0.82 Sat 492 407 479 534 619 622 534 461 0.77 0.82 Ce 363 35.4 27.5 35.4 49.2 37.7 1.32 17.3 2.66 Nd 1680 14.41 1.92 14.53 17.61 14.41 13.25 10.23 10.23 10.23 10.23 10.23 10.23 10.23	Zn	53.4	60.9	62.4	49.0	42.2	46.7	44.3	51.2	49.3
Rb 90.3 41.3 45.1 54.5 53.3 35.9 49.9 40.9 42.6 Sr 622 622 688 757 626 768 780 739 719 Y 9.9 9.3 116 6.8 9.1 41 7.1 5.0 5.0 Zr 120 106 102 137 135 139 122 139 117 Nb 5.22 4.95 3.73 3.31 3.07 2.97 3.46 4.05 2.79 Cs 3.87 1.32 1.73 2.86 1.52 0.82 1.41 0.77 0.82 La 21.5 18.3 16.2 2.2 2.65 2.12 1.66 4.60 1.49 1.55 1.62 1.63 La 21.5 1.50 1.51 1.44 1.32 2.17 1.62 1.63 Sm 3.03 0.24 0.79 0.68 0.68<	Ga	18.7	17.3	18.5	17.9	16.0	17.8	18.2	19.6	17.7
Sr 622 622 688 757 626 768 780 738 719 Y 99 9.3 116 6.8 9.1 4.1 7.1 5.0 5.0 Zr 120 106 102 137 135 139 122 139 17 Nb 5.22 4.95 3.73 3.31 3.07 2.97 3.46 4.05 2.79 Cs 3.87 1.32 1.73 2.86 1.52 0.82 1.41 0.77 0.82 Ba 492 407 479 534 619 462 534 461 600 La 21.5 18.3 16.2 2.22 26.5 1.22 19.6 14.9 13.5 Ce 36.3 3.5.4 27.5 3.5.4 49.2 3.7 3.3 2.73 2.17 Nd 16.80 14.41 13.92 14.53 17.61 14.41 13.25 10.23 9.44 Sm 3.03 2.51 2.68 2.09 2.61 2.32 2.17 1.62 1.65 Gd 2.30 2.02 2.15 1.50 1.91 1.56 1.66 </td <td>Rb</td> <td>90.3</td> <td>41.3</td> <td>45.1</td> <td>54.5</td> <td>53.3</td> <td>35.9</td> <td>49.9</td> <td>40.9</td> <td>42.6</td>	Rb	90.3	41.3	45.1	54.5	53.3	35.9	49.9	40.9	42.6
Y 9.9 9.3 11.6 6.8 9.1 4.1 7.1 5.0 5.0 Zr 120 106 102 137 135 139 122 139 17 Nb 5.22 4.95 3.73 3.31 3.07 2.97 3.46 4.05 2.79 Ba 492 407 479 534 619 462 534 461 600 La 21.5 18.3 16.2 2.2 2.65 21.2 19.6 14.9 13.5 Ce 36.3 3.5.4 2.75 3.54 4.92 3.3.7 3.13 2.73 2.17 Nd 1680 14.41 13.92 14.53 17.61 14.41 13.25 10.23 9.44 Sm 3.03 2.51 2.68 2.09 2.61 2.23 2.17 1.62 1.63 Gd 2.30 2.02 2.15 1.50 1.91 1.56	Sr	622	622	688	757	626	768	780	738	719
L7 L20 106 102 137 135 199 L22 199 117 Nb 5.22 4.95 3.73 3.31 3.07 2.97 3.46 4.05 2.79 Cs 3.87 1.32 1.73 2.86 1.52 0.82 1.41 0.77 0.82 La 21.5 18.3 16.2 2.2.2 26.5 21.2 19.6 14.9 13.5 Ce 36.3 35.4 27.5 35.4 49.2 33.7 31.3 27.3 21.7 Nd 16.80 14.41 13.92 14.53 7.61 14.41 13.25 10.23 9.44 Sm 3.03 2.51 2.68 2.09 2.61 2.23 2.17 1.66 1.27 1.31 Sm 3.03 0.28 0.33 0.17 0.27 0.19 0.23 0.15 0.18 Qu 1.69 1.54 1.88 0.85 <	Y	9.9	9.3	11.6	6.8	9.1	4.1	7.1	5.0	5.0
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	Zľ	120	106	102	13/	135	139	122	139	11/
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	IND Cs	3.22	4.95	5.75 1.73	2.86	3.07	2.97	5.40 1.41	4.05	2.79
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	Ba	492	407	479	534	619	462	534	461	600
Ce 36.3 35.4 27.5 35.4 49.2 33.7 31.3 27.3 21.7 Pr 4.53 3.95 3.60 4.17 5.04 4.06 3.77 2.90 2.61 Sm 3.03 2.51 2.68 2.09 2.61 2.23 2.17 1.62 1.63 Eu 0.81 0.81 0.79 0.68 0.68 0.71 0.67 0.52 0.66 Gd 2.30 2.02 2.15 1.50 1.91 1.56 1.66 1.27 1.31 Tb 0.33 0.28 0.33 0.17 0.27 0.19 0.23 0.15 0.18 Dy 1.69 1.54 1.88 0.85 1.31 0.81 0.22 0.16 0.61 Er 0.86 0.87 1.05 0.63 0.84 0.38 0.62 0.45 0.43 Im 0.13 0.13 0.17 0.11 0.14 0.52	La	21.5	18.3	16.2	22.2	26.5	21.2	19.6	14.9	13.5
Pr4.533.953.604.175.044.063.772.902.61Nd16.8014.4113.9214.5317.6114.4113.2510.239.44Sm3.032.512.682.092.612.232.171.621.63Eu0.810.810.790.680.680.710.670.520.66Gd2.302.022.151.501.911.561.661.271.31Dy1.691.541.880.851.310.811.200.820.89Ho0.310.300.360.180.270.140.220.160.16Er0.860.871.010.110.140.050.090.060.06Yb0.790.941.110.720.950.360.590.450.38Lu0.130.150.180.110.160.060.100.070.06Hf3.232.982.873.483.443.433.123.632.89Ta0.420.360.330.270.240.140.200.180.10Pb13.712.416.216.517.814.918.716.914.5Ta0.420.360.330.270.240.140.200.180.10Lu17.90.871.600.546.694.034.212	Ce	36.3	35.4	27.5	35.4	49.2	33.7	31.3	27.3	21.7
Nd16.8014.4113.9214.5317.6114.4113.2510.239.44Sm3.032.512.682.092.612.232.171.621.63Eu0.810.810.790.680.680.710.670.520.66Gd2.302.022.151.501.911.561.661.271.31Tb0.330.280.330.170.270.190.230.150.18Ho0.310.300.360.180.270.140.220.160.16Er0.860.871.050.630.840.380.620.450.43Tm0.130.130.170.110.140.050.990.450.38Lu0.130.150.180.110.160.060.100.070.06Yb0.790.941.110.720.950.360.590.450.38Lu0.130.150.180.110.160.060.100.070.06Yb0.790.941.110.720.950.360.590.450.38Lu0.130.150.180.110.160.060.100.070.06Yb0.790.941.110.720.240.140.200.180.10Lu1.371.2.416.216.517.814.918.7	Pr	4.53	3.95	3.60	4.17	5.04	4.06	3.77	2.90	2.61
Sm3.032.512.682.092.612.232.171.621.63Eu0.810.810.790.680.680.710.670.520.66Gd2.302.022.151.501.911.561.661.271.31Tb0.330.280.330.170.270.190.230.150.18Dy1.691.541.880.851.310.811.200.820.89Ho0.310.300.360.180.270.140.220.160.16Er0.860.871.050.630.840.380.620.450.43Tm0.130.130.170.110.140.050.090.060.06Vb0.790.941.110.720.950.360.590.450.38Lu0.130.150.180.110.160.060.100.070.06Hf3.232.982.873.483.443.433.123.632.89Ta0.420.360.330.270.240.140.200.180.10U1.790.871.600.540.830.380.570.410.38Sr/Y62.766.959.5111.869.0189.4110.1148.3143.8(La/Yb),19.51.4010.421.920.042.123.7<	Nd	16.80	14.41	13.92	14.53	17.61	14.41	13.25	10.23	9.44
Eu 0.81 0.79 0.68 0.68 0.71 0.67 0.52 0.63 Gd 2.30 2.02 2.15 1.50 1.91 1.56 1.66 1.27 1.31 Dy 1.69 1.54 1.88 0.85 1.31 0.81 1.20 0.82 0.89 Ho 0.31 0.30 0.36 0.18 0.27 0.14 0.22 0.16 0.16 Er 0.86 0.87 1.05 0.63 0.84 0.38 0.62 0.45 0.43 Tm 0.13 0.13 0.17 0.11 0.14 0.05 0.09 0.06 0.06 Lu 0.13 0.15 0.18 0.11 0.16 0.06 0.10 0.07 0.06 Hf 3.23 2.98 2.87 3.48 3.44 3.43 3.12 3.63 2.89 Ta 0.42 0.36 0.33 0.27 0.24 0.14	Sm	3.03	2.51	2.68	2.09	2.61	2.23	2.17	1.62	1.63
Gd2.302.022.151.501.911.561.661.271.31Tb0.330.280.330.170.270.190.230.150.18Dy1.691.541.880.851.310.811.200.820.89Ho0.310.300.360.180.270.140.220.160.16Er0.860.871.050.630.840.380.620.450.43Tm0.130.130.170.110.140.050.090.060.06Yb0.790.941.110.720.950.360.590.450.38Lu0.130.150.180.110.160.060.100.070.06Hf3.232.982.873.483.443.433.123.632.89Lu0.130.150.180.110.160.060.100.070.06Hf3.232.982.873.483.443.433.123.632.89Lu0.130.151.600.540.830.380.570.410.38Sr/Y6.2766.959.5111.869.0189.4110.1148.3143.8La/Yb/N19.51.4010.421.920.042.123.723.725.18^7Sr/8^6Sr0.7062390.7047200.704500.7046930.704525 <td>Eu</td> <td>0.81</td> <td>0.81</td> <td>0.79</td> <td>0.68</td> <td>0.68</td> <td>0.71</td> <td>0.67</td> <td>0.52</td> <td>0.66</td>	Eu	0.81	0.81	0.79	0.68	0.68	0.71	0.67	0.52	0.66
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Ga	2.30	2.02	2.15	1.50	1.91	1.56	1.66	1.27	1.31
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	DV	1.55	1.54	1.88	0.17	0.27	0.19	1.20	0.15	0.18
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Но	0.31	0.30	0.36	0.85	0.27	0.14	0.22	0.02	0.05
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Er	0.86	0.87	1.05	0.63	0.84	0.38	0.62	0.45	0.43
Yb 0.79 0.94 1.11 0.72 0.95 0.36 0.59 0.45 0.38 Lu 0.13 0.15 0.18 0.11 0.16 0.06 0.10 0.07 0.06 Hf 3.23 2.98 2.87 3.48 3.44 3.43 3.12 3.63 2.89 Ta 0.42 0.36 0.33 0.27 0.24 0.14 0.20 0.18 0.10 Pb 13.7 12.4 16.2 16.5 17.8 14.9 18.7 16.9 14.5 Th 6.95 3.45 5.29 4.06 6.69 4.03 4.21 2.38 1.84 U 1.79 0.87 1.60 0.54 0.83 0.38 0.57 0.41 0.38 Sr/Y 62.7 66.9 59.5 111.8 69.0 189.4 110.1 148.3 143.8 (La/Yb)_N 19.5 14.0 10.4 21.9 20.0 42.1 23.7 23.7 25.1 $^{87}Sr/^{86}Sr_i$ 0.70458 0.70459 0.704700 0.704693 0.704555 0.704565 0.704562 0.70454 $^{78}Sr/^{86}Sr_i$ 0.7048 0.7044 0.7044 0.7044 0.7044 0.7044 0.7044 $^{143}Nd/^{144}Nd$ 0.512637 0.51264 0.512664 0.512707 0.512673 0.512702 0.51265 $^{143}Nd/^{144}Nd$ 0.512637 0.51265 0.51265 0.51265 0.51265 <	Tm	0.13	0.13	0.17	0.11	0.14	0.05	0.09	0.06	0.06
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Yb	0.79	0.94	1.11	0.72	0.95	0.36	0.59	0.45	0.38
Hf 3.23 2.98 2.87 3.48 3.44 3.43 3.12 3.63 2.89 Ta 0.42 0.36 0.33 0.27 0.24 0.14 0.20 0.18 0.10 Pb 13.7 12.4 16.2 16.5 17.8 14.9 18.7 16.9 14.5 Th 6.95 3.45 5.29 4.06 6.69 4.03 4.21 2.38 1.84 U 1.79 0.87 1.60 0.54 0.83 0.38 0.57 0.41 0.38 Sr/Y 62.7 66.9 59.5 111.8 69.0 189.4 110.1 148.3 143.8 (La/Yb)_N 19.5 14.0 10.4 21.9 20.0 42.1 23.7 23.7 25.1 $^{87}Sr/^{86}Sr$ 0.704529 0.704720 0.704700 0.704693 0.704525 0.704562 0.704561 $^{87}Sr/^{86}Sr_i$ 0.7045 0.7044 0.7044 0.7044 0.7044 0.7044 0.7044 0.7044 $^{143}Nd/^{144}Nd_i$ 0.512637 0.512699 0.51264 0.512747 0.512673 0.512702 0.512709 $^{143}Nd/^{144}Nd_i$ 0.51258 0.51264 0.51266 0.51267 0.51265 0.51265 0.51265 $e_Nd(T)$ 0.9 2.2 3.0 1.6 3.2 2.4 1.7 2.3 2.3 $T_{DM}(Ga)$ 0.75 0.64 0.63 0.59 0.51 0.56	Lu	0.13	0.15	0.18	0.11	0.16	0.06	0.10	0.07	0.06
Ta 0.42 0.36 0.33 0.27 0.24 0.14 0.20 0.18 0.10 Pb 13.7 12.4 16.2 16.5 17.8 14.9 18.7 16.9 14.5 Th 6.95 3.45 5.29 4.06 6.69 4.03 4.21 2.38 1.84 U 1.79 0.87 1.60 0.54 0.83 0.38 0.57 0.41 0.38 Sr/Y 62.7 66.9 59.5 111.8 69.0 189.4 110.1 148.3 143.8 (La/Yb)_N 19.5 14.0 10.4 21.9 20.0 42.1 23.7 23.7 25.1 $^{87}Sr/^{86}Sr$ 0.704529 0.704720 0.704591 0.704693 0.704525 0.704655 0.704562 0.704581 $^{87}Sr/^{86}Sr_i$ 0.7045 0.7044 0.7044 0.7044 0.7044 0.7044 0.7044 0.7044 $^{143}Nd/^{144}Nd_i$ 0.512637 0.512699 0.51264 0.512747 0.512673 0.512702 0.512709 $^{143}Nd/^{144}Nd_i$ 0.51258 0.51264 0.51266 0.51266 0.512673 0.51265 0.51265 $\epsilon_{Nd}(T)$ 0.9 2.2 3.0 1.6 3.2 2.4 1.7 2.3 2.3 $T_{DM}(Ga)$ 0.75 0.64 0.63 0.59 0.51 0.56 0.64 0.58 0.62 T_{C}^{-0} 741 730 724 760 <	Hf	3.23	2.98	2.87	3.48	3.44	3.43	3.12	3.63	2.89
Pb13.712.416.216.517.814.918.716.914.5Th6.953.455.294.066.694.034.212.381.84U1.790.871.600.540.830.380.570.410.38Sr/Y62.766.959.5111.869.0189.4110.1148.3143.8(La/Yb)N19.514.010.421.920.042.123.723.725.1 $^{87}Sr/^{86}Sr0.7052390.7047200.7045910.7047000.7046930.7045250.7046550.7045620.704581^{87}Sr/^{86}Sr_i0.70480.70450.70440.70440.70440.70440.70440.7044^{143}Nd/^{144}Nd_i0.5126370.5126990.5127480.5126640.5127470.5127060.5126730.5127020.512709^{143}Nd/^{144}Nd_i0.512580.512640.512620.512650.512650.512650.512650.51265\epsilon_{Nd}(T)0.92.23.01.63.22.41.72.32.3T_{DM}(Ga)0.750.640.630.590.510.560.640.580.62T_{X'}^{eC}(C)741730724760761758752755744$	Ta	0.42	0.36	0.33	0.27	0.24	0.14	0.20	0.18	0.10
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	PD Th	13.7	12.4	16.2	16.5	17.8	14.9	18.7	16.9	14.5
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	III II	0.95	5.45 0.87	1.60	4.00	0.09	4.05	4.21	2.58	0.38
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Sr/Y	62.7	66.9	59.5	111.8	69.0	189.4	110.1	148.3	143.8
	(La/Yb) _N	19.5	14.0	10.4	21.9	20.0	42.1	23.7	23.7	25.1
	⁸⁷ Sr/ ⁸⁶ Sr	0.705239	0.704720	0.704591	0.704700	0.704693	0.704525	0.704655	0.704562	0.704581
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	⁸⁷ Sr/ ⁸⁶ Sr _i	0.7048	0.7045	0.7044	0.7045	0.7044	0.7044	0.7044	0.7044	0.7044
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	¹⁴³ Nd/ ¹⁴⁴ Nd	0.512637	0.512699	0.512748	0.512664	0.512747	0.512706	0.512673	0.512702	0.512709
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	¹⁴³ Nd/ ¹⁴⁴ Nd i	0.51258	0.51264	0.51269	0.51262	0.51270	0.51266	0.51262	0.51265	0.51265
$T_{DM}(Ga)$ 0.75 0.64 0.63 0.59 0.51 0.56 0.64 0.58 0.62 $T_{rd}^{\circ}C$) 741 730 724 760 761 758 752 765 744	$\varepsilon_{\rm Nd}(1)$	0.9	2.2	3.0	1.6	3.2	2.4	1./	2.3	2.3
	T_{DM} (Gd) $T_{T_{a}}$ (°C)	741	730	724	760	761	758	752	765	744

*: total Fe, A/CNK: molar Al₂O₃/(CaO+Na₂O+K₂O), Mg#=100(Mg/Mg+Fe), n.d.: not determined. *⁷Sr/⁸⁶Sr₁=⁸⁷Sr/⁸⁶Sr⁻(⁸⁷Rb/⁸⁶Sr)×(e^X⁻-1), λ_{Rb-Sr} =0.0142 Ga⁻¹, ⁸⁷Rb/⁸⁶Sr=(Rb/Sr)×2.8956. ¹⁴³Nd/¹⁴⁴Nd₁=¹⁴³Nd/¹⁴⁴Nd⁻(¹⁴⁷Sm/¹⁴⁴Nd)×(e^{XT}-1), λ_{Sm-Nd} =0.00654 Ga⁻¹, ¹⁴⁷Sm/¹⁴⁴Nd=(Sm/Nd)×0.60456. $\varepsilon_{Nd}(T) = [(^{143}Nd/^{144}Nd)_{Sample}(T)/(^{143}Nd/^{144}Nd)_{CHUR}(T)-1]\times10^4$, (¹⁴³Nd/¹⁴⁴Nd)_{CHUR}(T)=0.512638-0.1967×(ε^{NT} -1). T_{DM} (Ga)=1/ λ_{Sm-Nd} ×In {1+[((¹⁴³Nd/¹⁴⁴Nd)_{Sample}-0.51315)/((¹⁴⁷Sm/¹⁴⁴Nd)_{Sample}-0.2137)]} Samples not dated are calculated with an age (T) of 80 Ma.

 $T_{Zr}(^{\circ}C)$: zircon saturation temperature calculated by using the Watson and Harrison (1983) thermometer.



Fig. 5. A/CNK [molar $Al_2O_3/(CaO + Na_2O + K_2O)$] vs. SiO₂ diagram. Comparison data of the Gangdese batholithic rocks are from Wen (2007).

Gangdese adakites, yielding temperatures from ~765 to 724 °C (Table 5) and pressures at $0.85\pm0.09(1\sigma)$ GPa (from ~1.1 to 0.66 GPa). These estimates suggest that the magmas were emplaced around ~30–25 km depths under which epidote and muscovite crystallized at ~750 °C near the solidus temperature.

The occurrence of subhedral epidote/muscovite crystals requires not only deep-seated emplacement but also a following rapid upward transport, otherwise the euhedral crystals cannot survive from complete dissolution by quartzo-feldspathic matrix if slow cooling took place (Zen, 1985; Brandon et al., 1996; Sial et al., 1999; Schmidt and Poli, 2004). Petrographic evidence supporting such a rapid upward transport, and fast cooling, is recorded as the disequilibrium texture (partial dissolution and resorption) surrounding the epidote, muscovite and plagioclase crystals (Fig. 3). This, as exemplified by the North American Cordillera (Ducea, 2001; Saleeby, 2003), may have been affiliated with tectonic uplift and exhumation that could expose mid-crustal (~25–30 km) batholithic rocks with a rate as fast as 2– 3 mm/yr during the early Laramide orogeny (Kidder et al., 2003).

6.4. Implications for flat subduction and regional tectonics

Based on zircon U–Pb age data, Wen et al. (in press) demonstrated that the Gangdese magmatism was active in the Late Cretaceous (ca. 103–80 Ma) and early Paleogene (ca. 68–43 Ma), with a magmatic gap or quiescent period that existed between ca. 80 and 68 Ma. It is furthermore proposed that, following a "normal-angle" subduction that resulted in the extensive Cretaceous Gangdese intrusions during ca. 100 and 85 Ma, a "flat-slab" stage of Neo-Tethyan subduction



Fig. 6. Chondrite-normalized REE variation diagram. Normalizing values are from Sun and McDonoud (1989). The Cretaceous Gangdese calc-alkaline rocks plotted for comparison are from Wen (2007).



Fig. 7. Primitive mantle-normalized trace element variation diagram. Normalizing values are from Sun and McDonoud (1989). The Cretaceous Gangdese calc-alkaline rocks plotted for comparison are from Wen (2007).

occurred so that led to a contractional tectonic regime lasting until the end-Cretaceous in southern Tibet. Exemplary analogues quoted are the Central Andes (Kay and Mpodozis, 2001) and the Late Cretaceous Laramide orogeny in southern North America (Saleeby, 2003), in both cases flattening of subducting slab has been postulated to not only tectonically thicken the arc crust but also squeeze out the mantle wedge, thus terminating the arc magmatism with adakites being generated in the ending phase (Kay and Mpodozis, 2001; Booker et al.,



Fig. 8. Plots of (a) Sr/Y vs. Y and (b) chondrite-normalized La/Yb ratios vs. Yb. Fields of adakite and TTG and arc calc-alkaline lavas are from Martin (1999). In (b) a Gangdese gabbro T036B (ca. 102 Ma; Wen et al., in press) was utilized as the source rock for the REE modeling under amphibolite and eclogite conditions, with various garnet contents and respective partition coefficients proposed by Irving and Frey (1978), Fujimaki et al. (1984) and Sisson (1994).



Fig. 9. Sr–Nd isotope correlation diagram of the Gangdese adakites. Field of the Neo-Tethyan ophiolites from the Yarlu Tsangpo suture representative of the subducted oceanic crust is constructed using the data reported by Xu and Castillo (2004) and Zhang et al. (2005). The upper crust (UC) and lower crust (LC) mixing curves are constructed using the following end-member parameters: (1) a Gangdese gabbro [sample T036B: $\varepsilon_{Nd}(T)$ =+3.7, ⁸⁷Sr/⁸⁶Sr_i=0.70435]; (2) UC: $\varepsilon_{Nd}(T)$ =-10.2, ⁸⁷Sr/⁸⁶Sr_i=0.7044, Nd=11 ppm, Sr=315 ppm; and (3) LC: $\varepsilon_{Nd}(T)$ =-10.2, ⁸⁷Sr/⁸⁶Sr_i=0.7044, Nd=11 ppm, Sr=348 ppm. All these and Gangdese data are from Wen (2007) except the Sr and Nd concentrations of the LC component are assumed values from Rudnick and Gao (2003).

2004). Mimicking scenarios are observed in southern Tibet where the Cretaceous stage of Gangdese intrusions ceased with adakitic activity followed by a magmatic gap. Wen et al. (in press), therefore, argued that the region underwent a combined magmatic/tectonic thickening of the crust above the mantle wedge, with the former owing to extensive basaltic underplating associated with the Cretaceous Gangdese intrusions and the latter to the proposed flat subduction. In the Andes, flat subduction occurs when oceanic plateau, ridge or seamount is being subducted (cf. Gutscher et al., 2000). Within the Neo-Tethys Ocean, similar structural features should have been common because their remnants have been widely identified along the Yarlu Tsangpo suture zone (Aitchison et al., 2004; Dubois-Cote et al., 2005; Dupuis et al., 2005).

This flat subduction model explains not only the magmatic evolution but also the Late Cretaceous crustal deformation and shortening in the Lhasa terrane (England and Searle, 1986; Coulon et al., 1986; Ratschbacher et al., 1993; Murphy et al., 1997; Yin and Harrison, 2000). The Cretaceous Takena Formation, for example, stopped deposition at ca. 90-80 Ma and experienced a period of folding and erosion from ca. 80 to 70 Ma, before the restart of nonmarine deposition accompanied by the Paleogene Linzizong volcanism (Leier et al., 2007). These can be interpreted as the consequence of an Andean- or "Pacific-type" accretionary orogeny (Maruyama, 1997) that happened in southern Tibet at this specific time span owing to the flat subduction. Analogous to the Central Andes that shows high surface elevations associated with thick crust and a welldeveloped fold-and-thrust belt in the retroarc region (cf. Kay and Mpodozis, 2001), it is rational to further argue that a high, but relatively narrow, mountain range existed in southern Tibet during the Late Cretaceous. In other words, southern Tibet was situated in an accretionary convergent margin where subduction-related orogenic processes involving arc magmatism, crustal thickening and rapid tectonic uplift had been operating actively prior to the collision of India with this particular part of Asia.

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